

Space TeraHertz Observatory (STO): A Terahertz Observatory for Origins Research

Christopher Walker¹, Craig Kulesa¹, Gordon Chin², David Fischer³,
Paul Goldsmith⁴, Mark Heyer⁵, David Hollenbach⁶, William Langer⁷,
Daniel Lester⁸, Imran Mehdi⁷, Gary Melnick⁹, Gopal Narayanan⁵,
Thomas Phillips¹⁰, Gordon Stacey⁴, Sander Weinreb⁷, Mark Wolfire¹¹,
Harold Yorke⁷, Erick Young¹, and Lucy Ziurys¹

(Email: cwalker@as.arizona.edu)

¹Department of Astronomy, University of Arizona, Tucson, Arizona

²NASA Goddard Space Flight Center, Greenbelt, Maryland

³Ball Aerospace

⁴Cornell University, Ithaca, New York

⁵Department of Astronomy, University of Massachusetts, Amherst, Massachusetts

⁶NASA Ames Research Center, Moffett Field, California

⁷Jet Propulsion Laboratory, Pasadena, California

⁸McDonald Observatory, University of Texas, Austin, Texas

⁹Smithsonian Astrophysical Observatory, Cambridge, Massachusetts

¹⁰California Institute of Technology, Pasadena, California

¹¹Department of Astronomy, University of Maryland, Colleger Park, Maryland

The 10-meter-class Space Terahertz Observatory (STO) has been proposed as a NASA Origins concept study and is designed to 1) conduct origin studies of planets, stars, and molecular clouds; 2) trace the life cycle of the Interstellar Medium (ISM) and star formation rate throughout the Galaxy; 3) measure the gas content of formative pre-planetary disks; and 4) observe the distribution of atomic and molecular gas in both nearby and distant galaxies. STO will achieve these goals through high spectral and angular resolution observations of C⁺, O, N⁺, HD, and H₂O lines in the far-infrared. The science goals of STO can be achieved either through a dedicated mission with an uncooled 8–10 meter primary or by implementing heterodyne instrumentation on SAFIR and extending its operational lifetime. As part of the NASA concept study, we will investigate the relative merits of these two approaches. The STO instrument concept can be used to define technological roadmaps for THz astronomy and instrument development in the coming decades. The instrument utilizes four 8 × 8 and two 4 × 4 heterodyne receiver arrays to produce a total of 288 diffraction-limited beams in the focal plane, yielding angular resolution from 1.7'' at 60 microns to 16'' at 540 microns. Each beam will produce a 2048-element spectrum with ~0.3 km/s resolution. STO science objectives are closely aligned with NASA's Origins Roadmap and drive the creation of a new generation of heterodyne array instrumentation that benefits directly from technologies developed for the Herschel HIFI instrument.